

Halo Occupation Distribution of Quasars

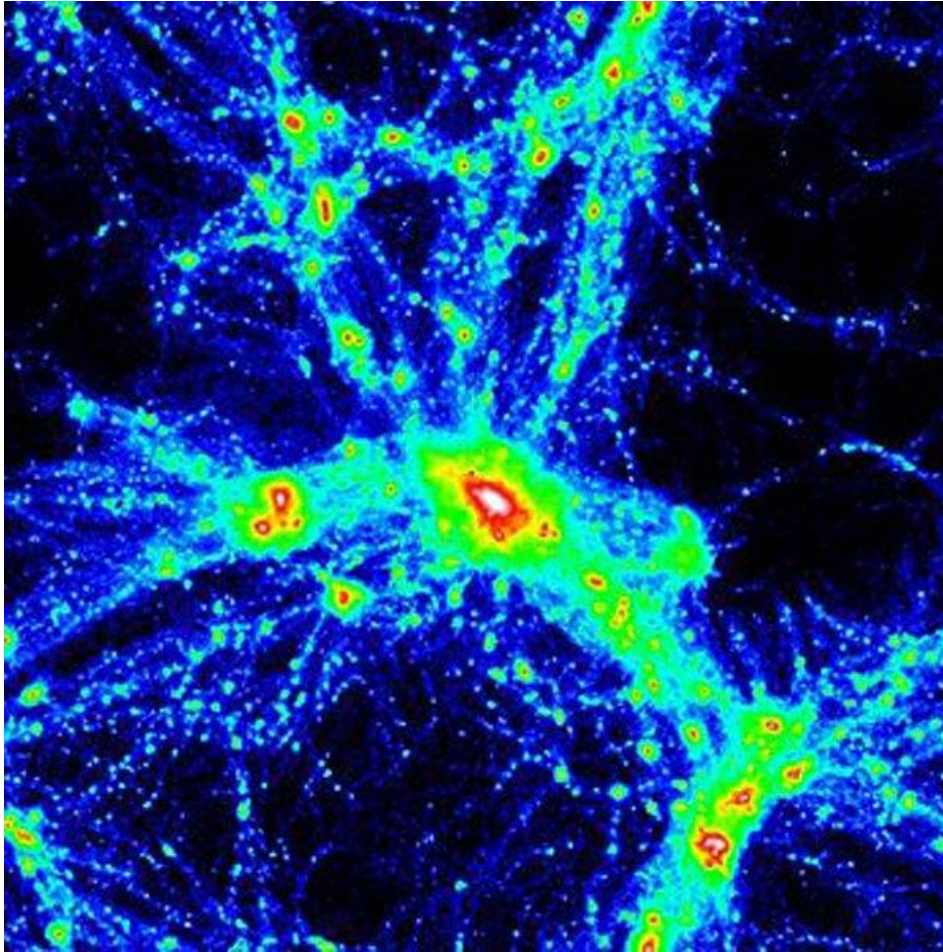
Kaustav Mitra

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Sep. 19, 2015

Dark Matter Halo

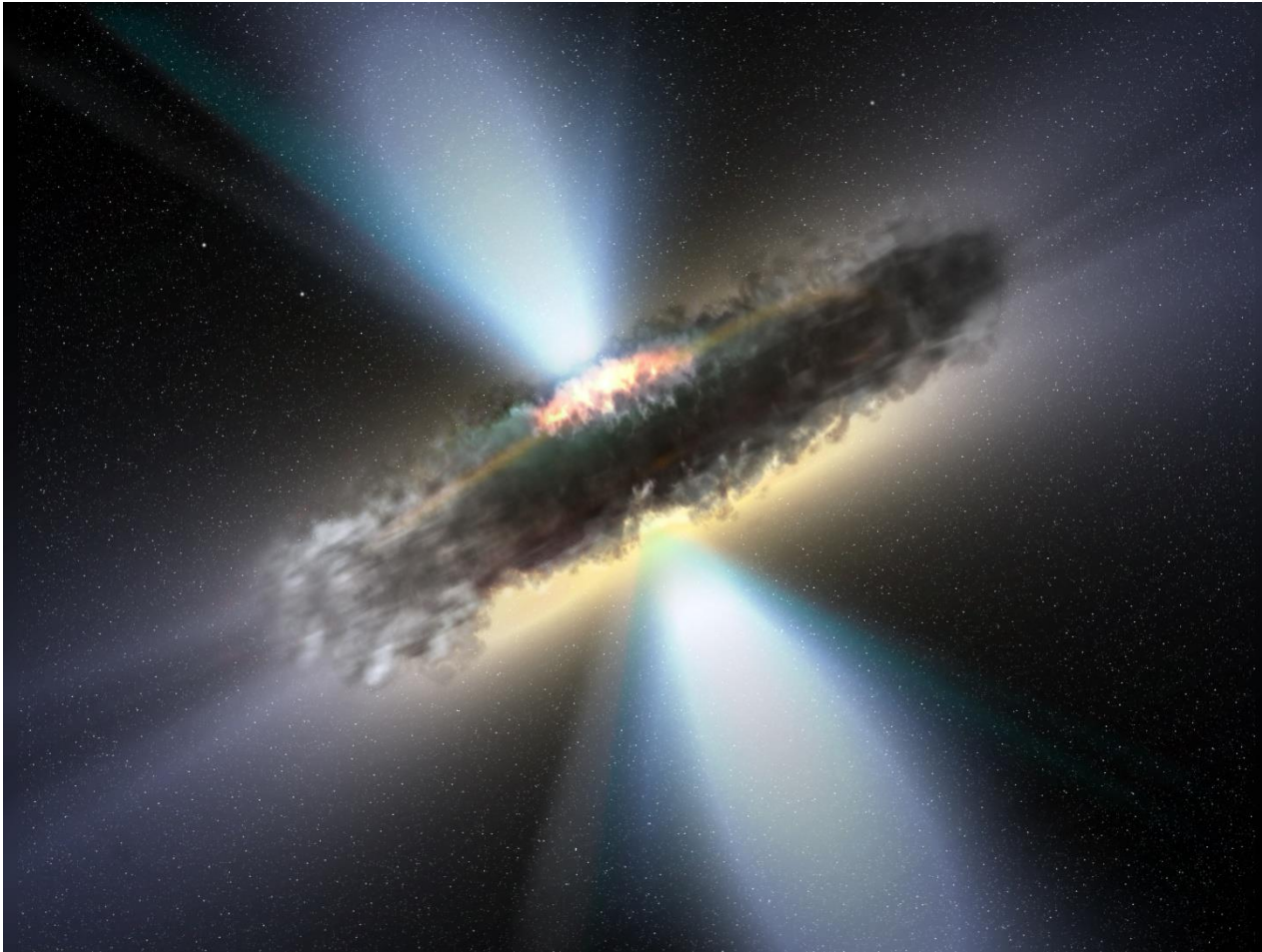


Galaxies lie in
dark matter
halos

Each galaxy
hosts a super
massive black
hole

32-Mpc simulation of large-scale structure in a standard cold-dark matter Universe.
Courtesy : R. Cen, Princeton University.

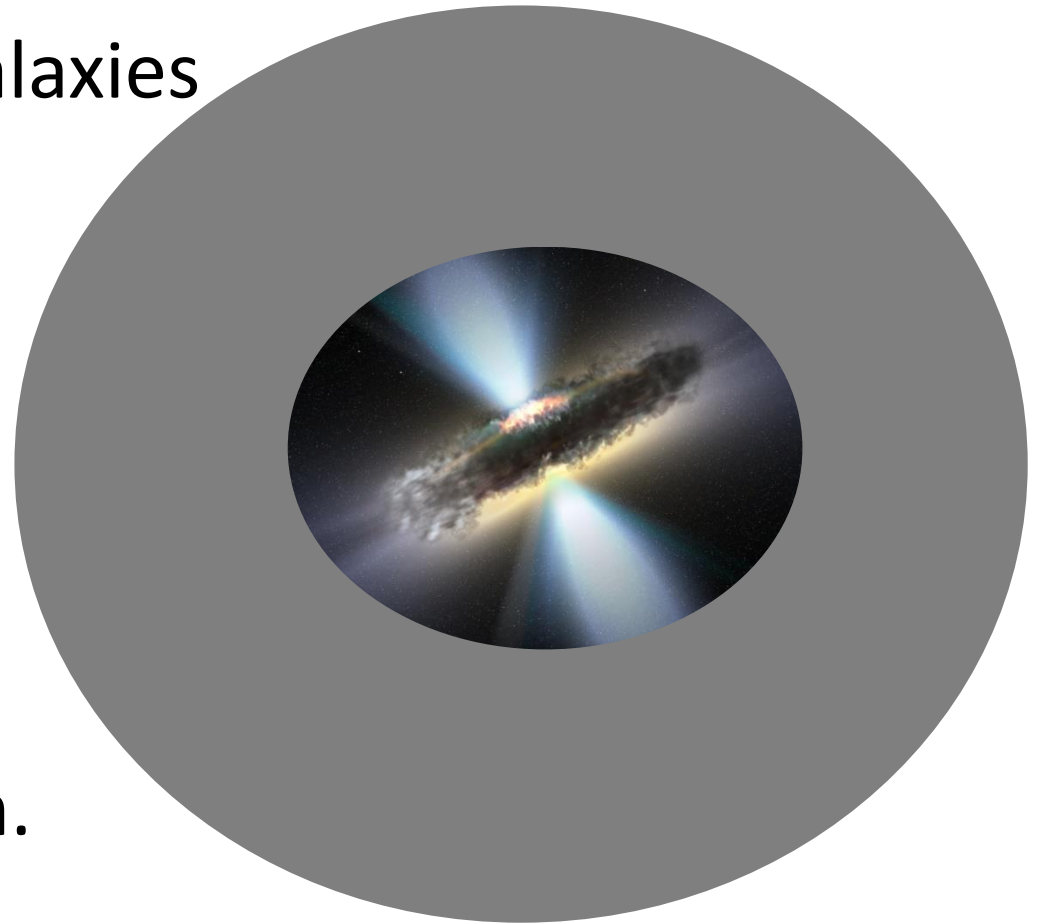
Quasars



Courtesy : ESA | V. Beckmann (NASA-GSFC)

Quasars in Dark Matter Halos

Co-evolution Of Galaxies
and Super Massive
Black Holes
with
Large
scale Dark
Matter distribution.



Clustering of Quasars



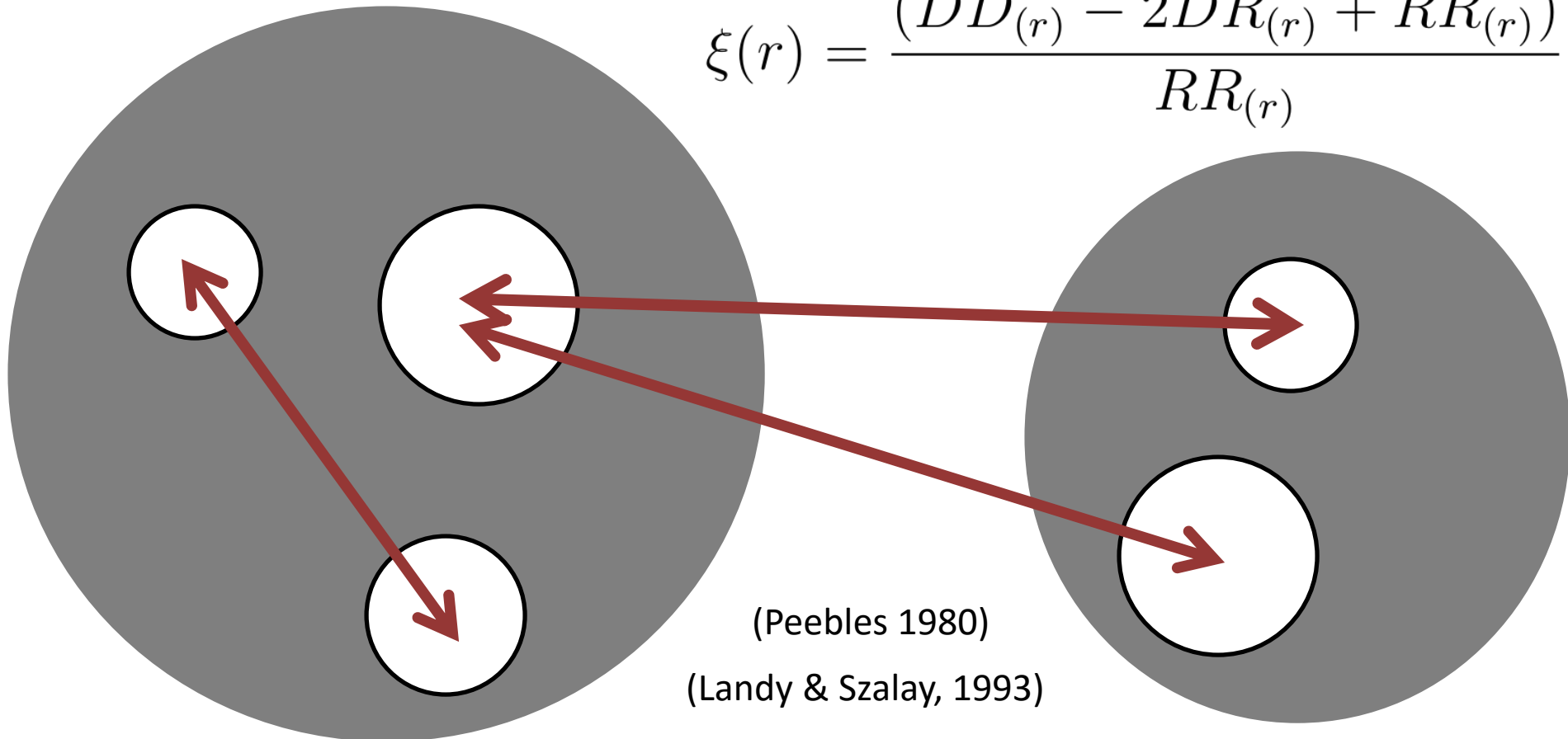
How are
quasars
distributed in
Dark Matter
halos?

The Coma Cluster. Credit: NASA, ESA, and the Hubble Heritage Team (STScI/AURA) .
Acknowledgment: M. Carter (Liverpool John Moores University) and the Coma HST ACS
Treasury Team.

Clustering Statistic

Counting the number of pairs of quasars at different clustering scales

$$\xi(r) = \frac{(DD_{(r)} - 2DR_{(r)} + RR_{(r)})}{RR_{(r)}}$$



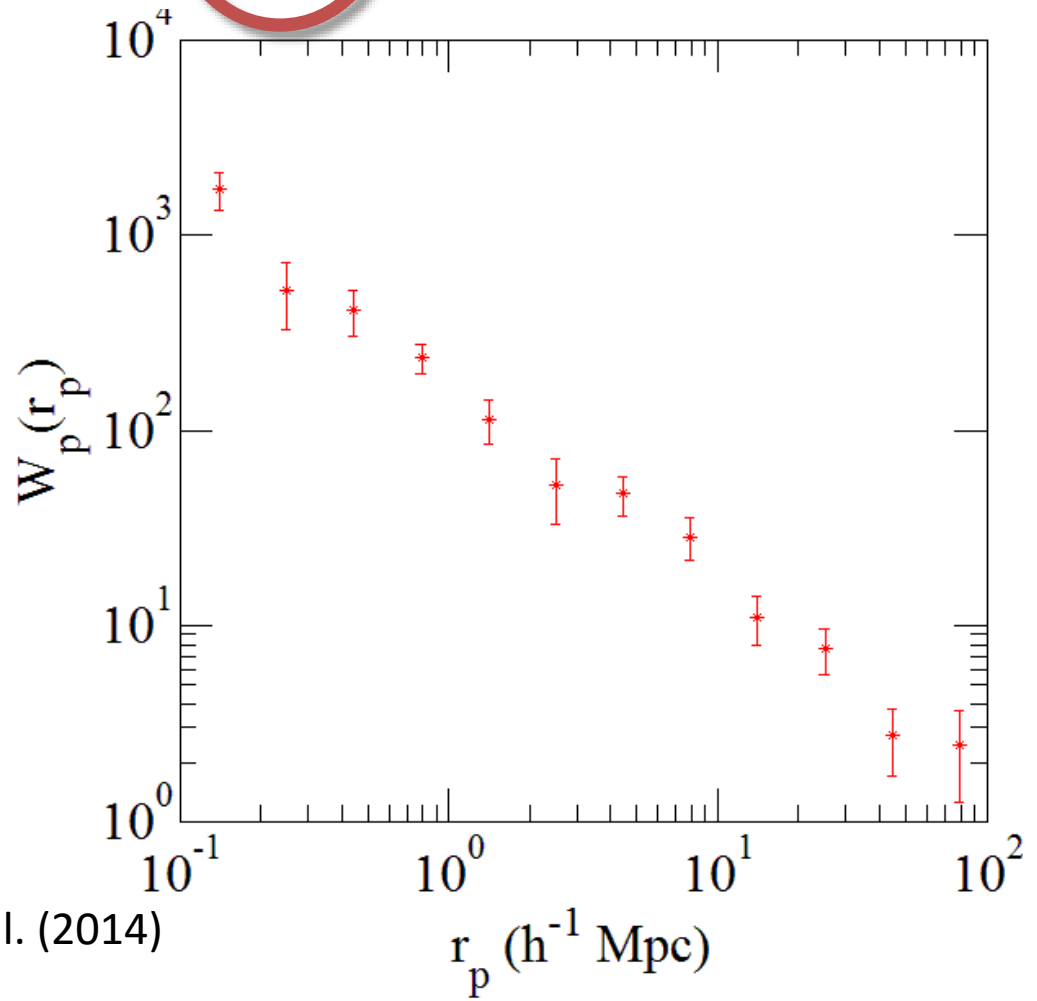
(Peebles 1980)

(Landy & Szalay, 1993)

Two Point Correlation Function

$$P(r) = n^2 [1 + \xi(r)] dV_1 dV_2 \quad (\text{Peebles 1980})$$

Excess probability
Of finding quasar
Pairs at that scale,
hence a measure
of clustering at
different clustering
scales.

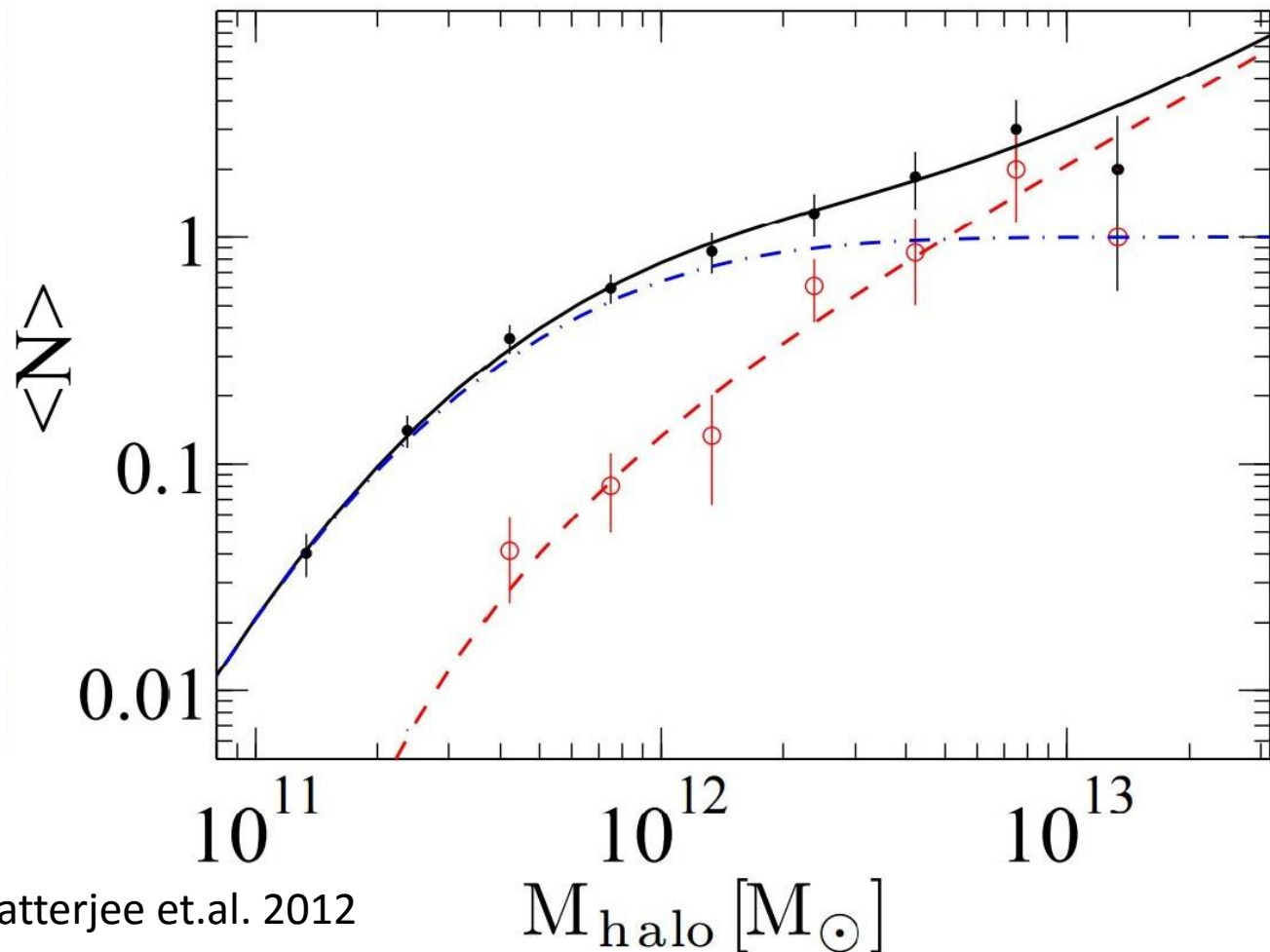


DiPompeo et al. (2014)

Mean Occupation Function

$$\langle N(M) \rangle = \frac{1}{2} \left[1 + \operatorname{erf} \left(\frac{\log M - \log M_{\min}}{\sigma_{\log M}} \right) \right] + \left(\frac{M}{M_1} \right)^\alpha \exp \left(-\frac{M_{\text{cut}}}{M} \right)$$

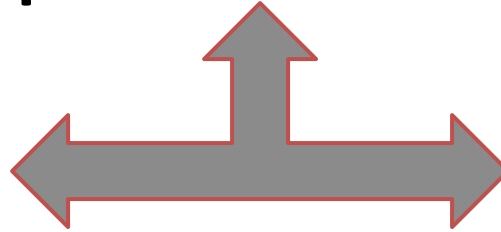
Average
number
of quasars
per dark
matter halo
of a given
host halo
mass.



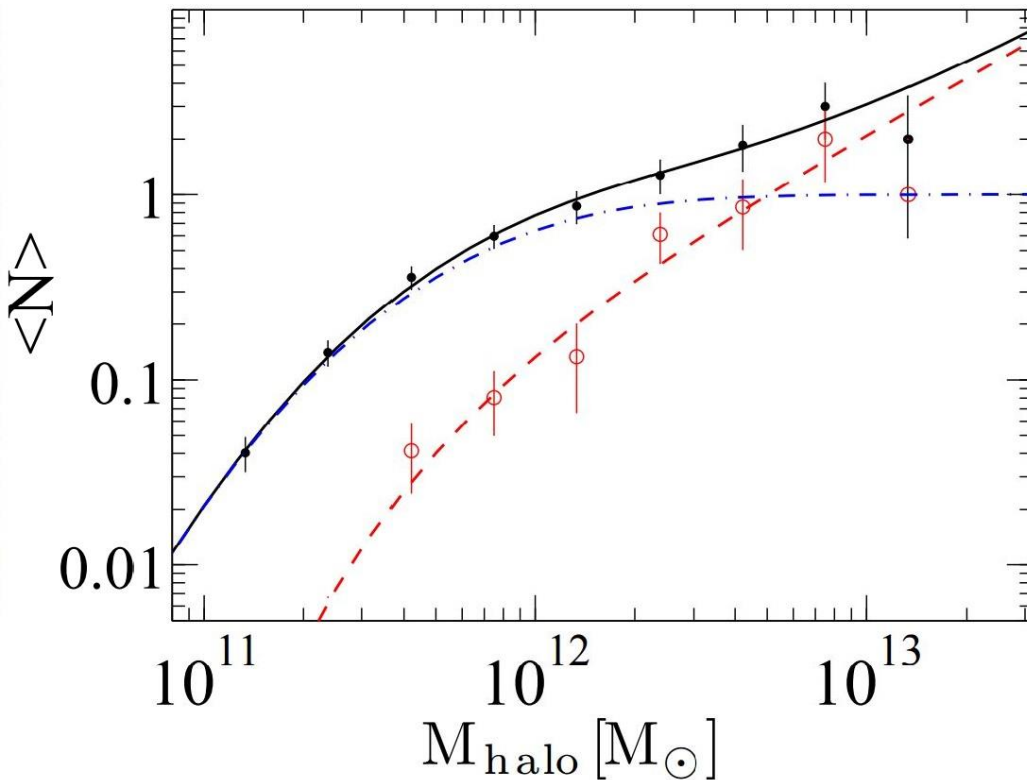
Chatterjee et.al. 2012

Halo Occupation Distribution

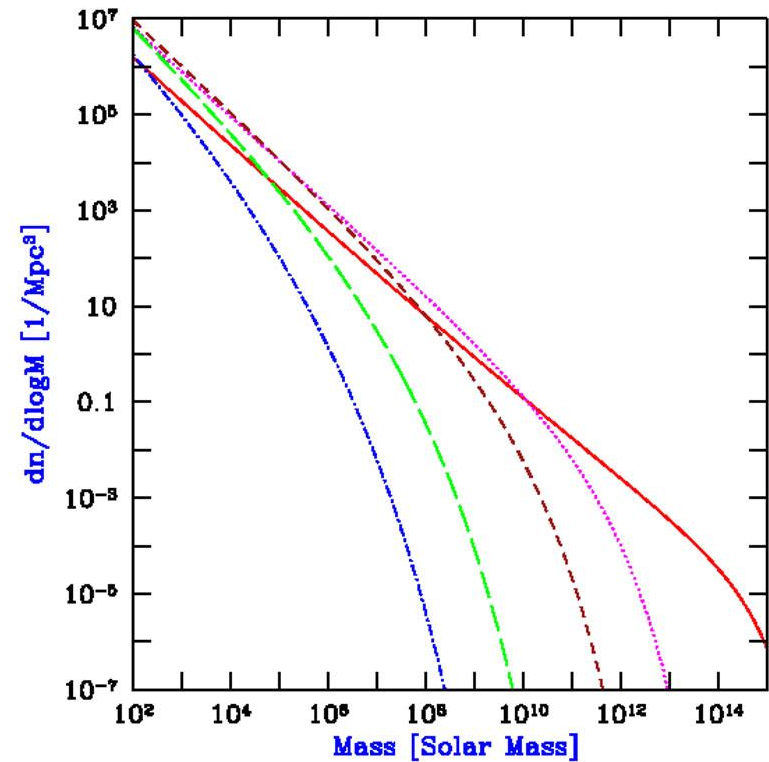
Mean Occupation
Function



Halo Mass
Function



Chatterjee et.al. 2012



Courtesy: Avi Loeb

So... what I did is as follows...

- Assumed Chatterjee et. al. 2012 MOF model and Tinker et. al. 2007 HMF.
- Used DiPompeo et. al. 2014 data of observed 2PCF
- Used Zheng et. al. 2007 MCMC code to find the set of free parameters that minimizes the chi-square of modeled 2PCF with that of the observed data.
- Used the free parameters to construct the **HOD of obscured and unobscured quasars.**

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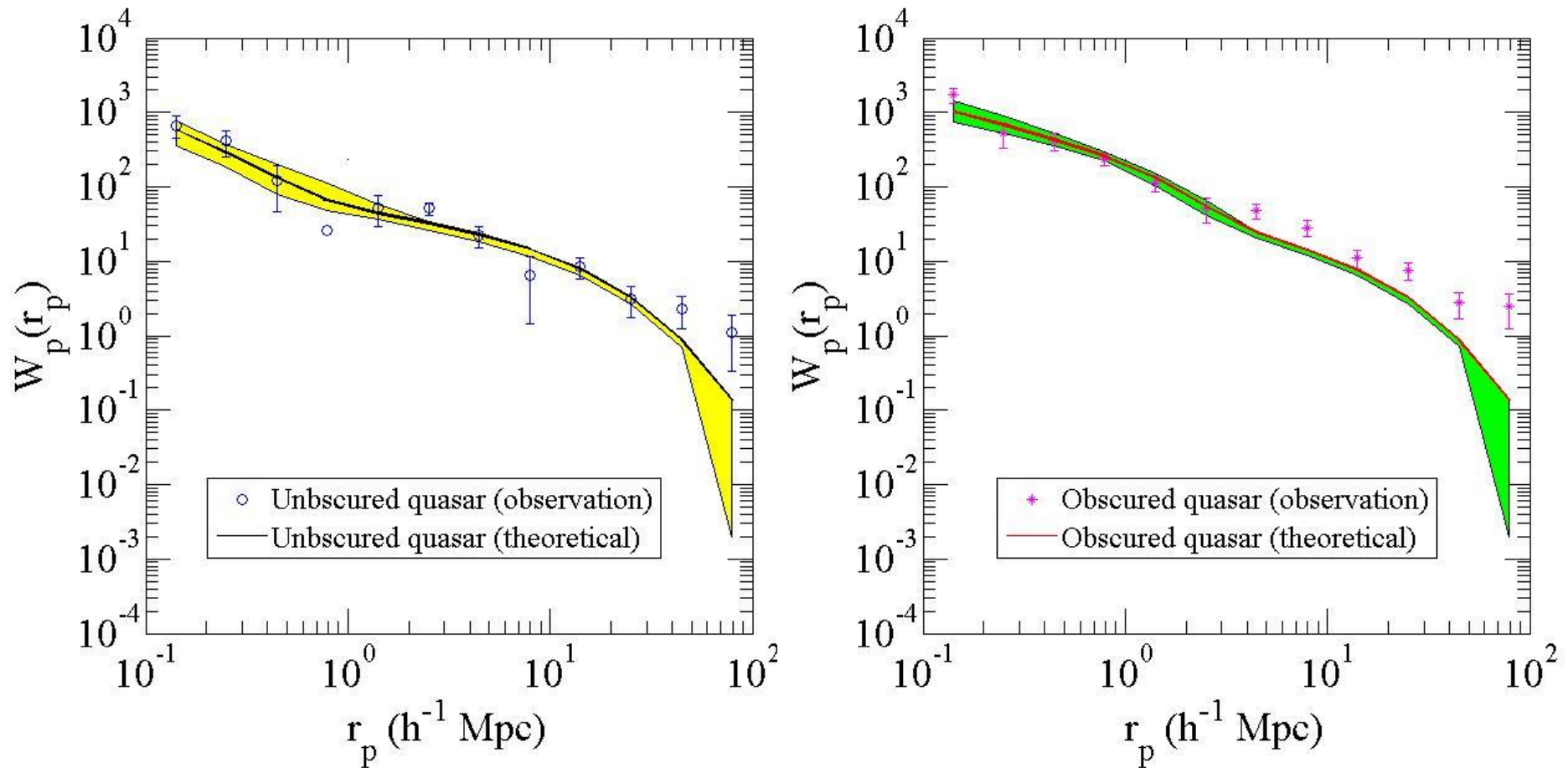
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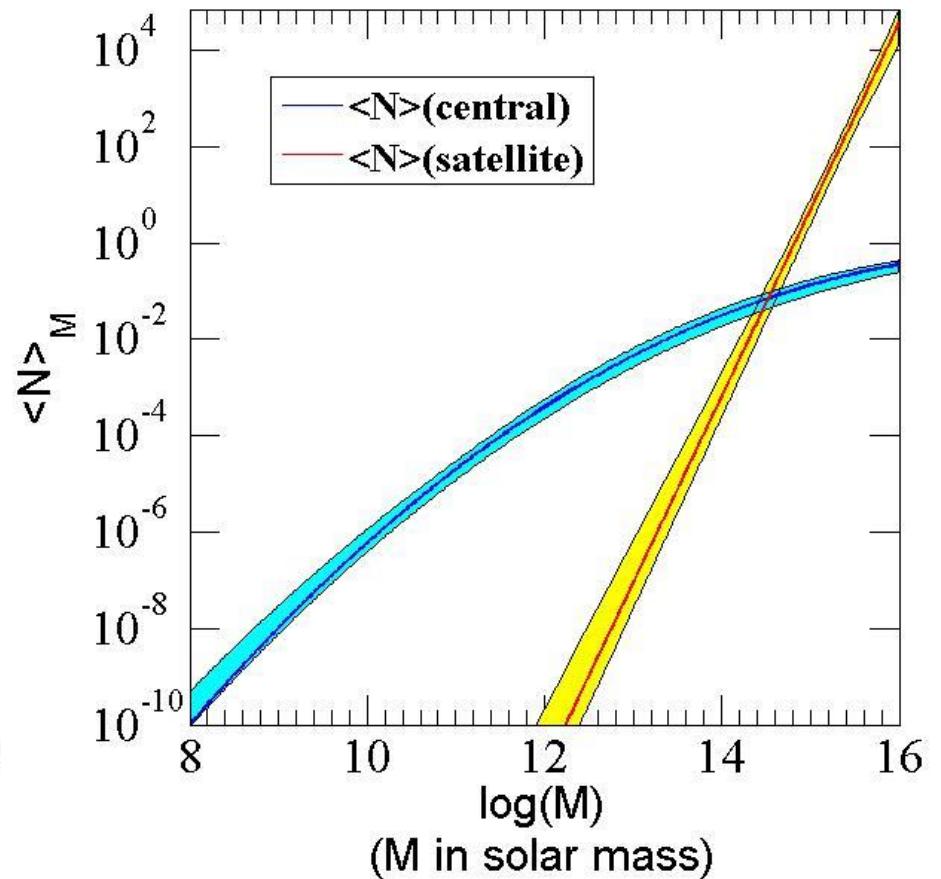
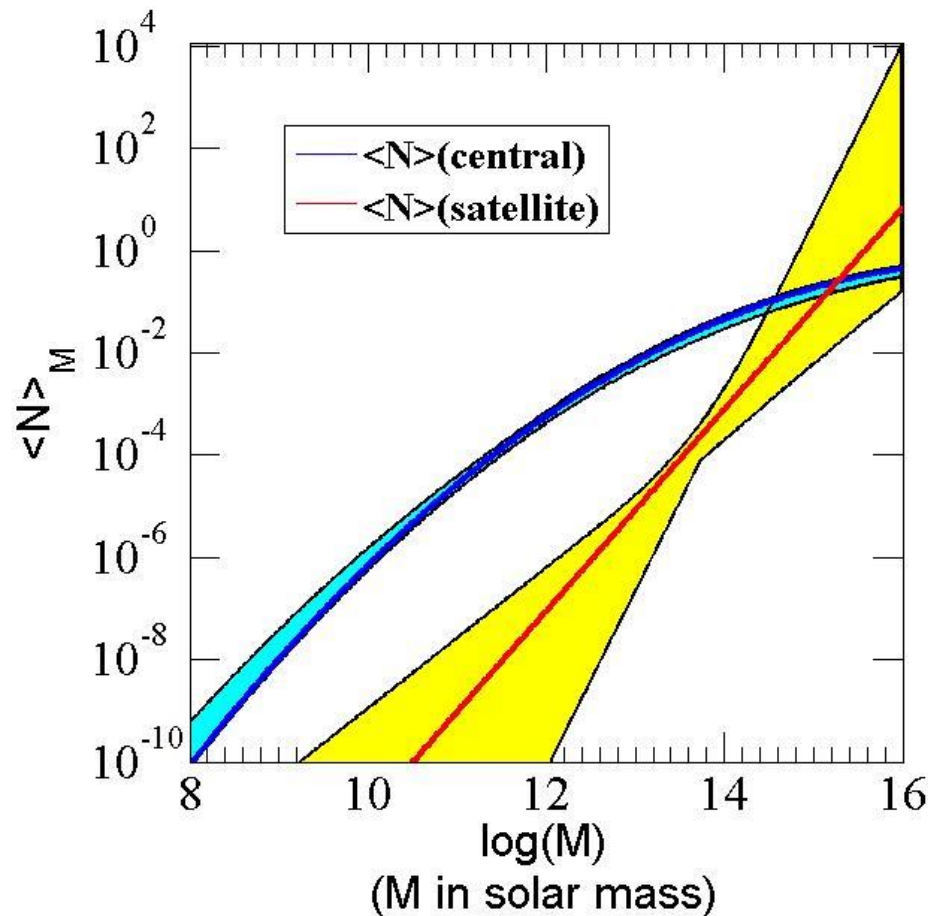
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Two Point Correlation Function of Obscured and Unobscured Quasars



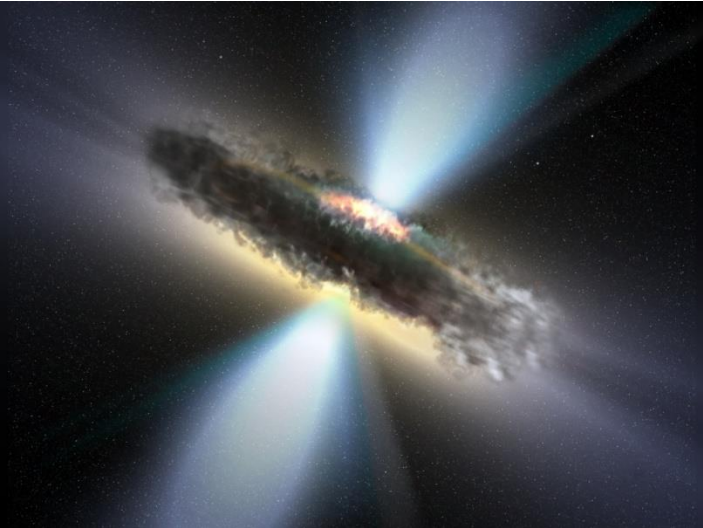
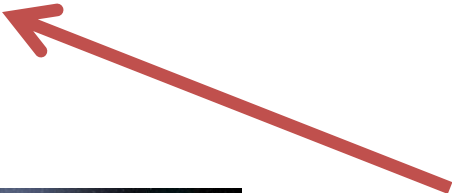
Data from DiPompeo et al. (2014)

Mean Occupation Function of Obscured and Unobscured Quasars

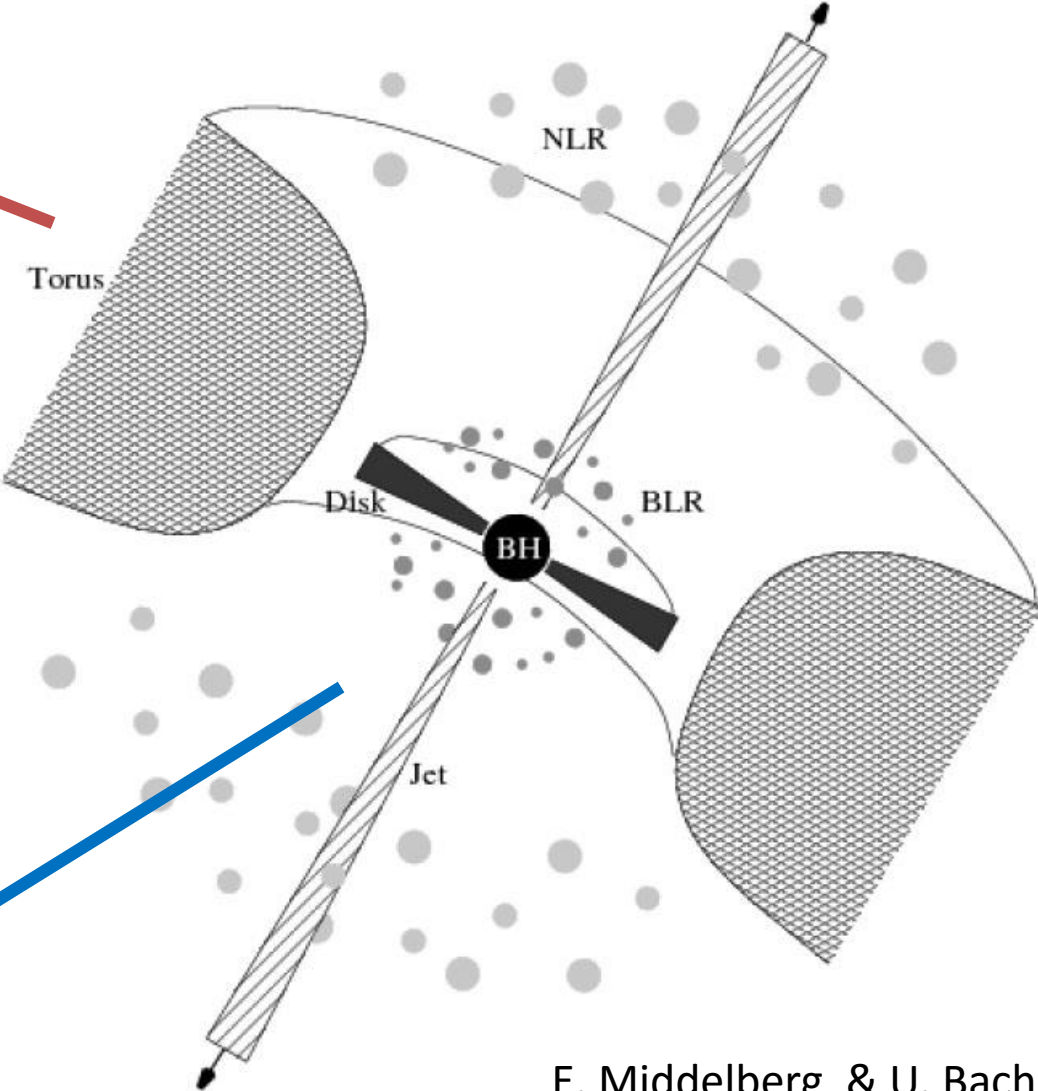


Quasars ...again

Obscured



ESA | V. Beckmann (NASA-GSFC)



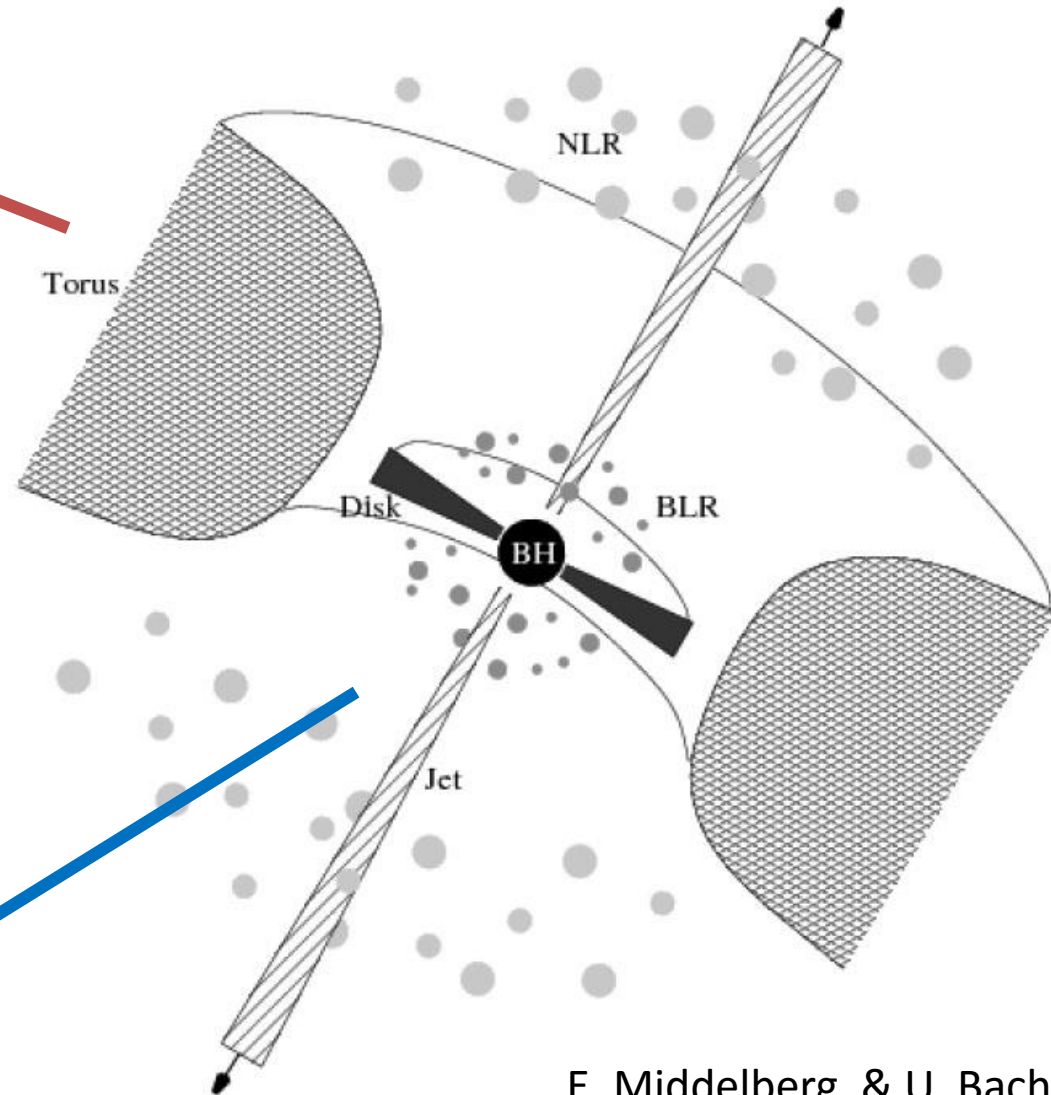
Unobscured

Are they intrinsically different?

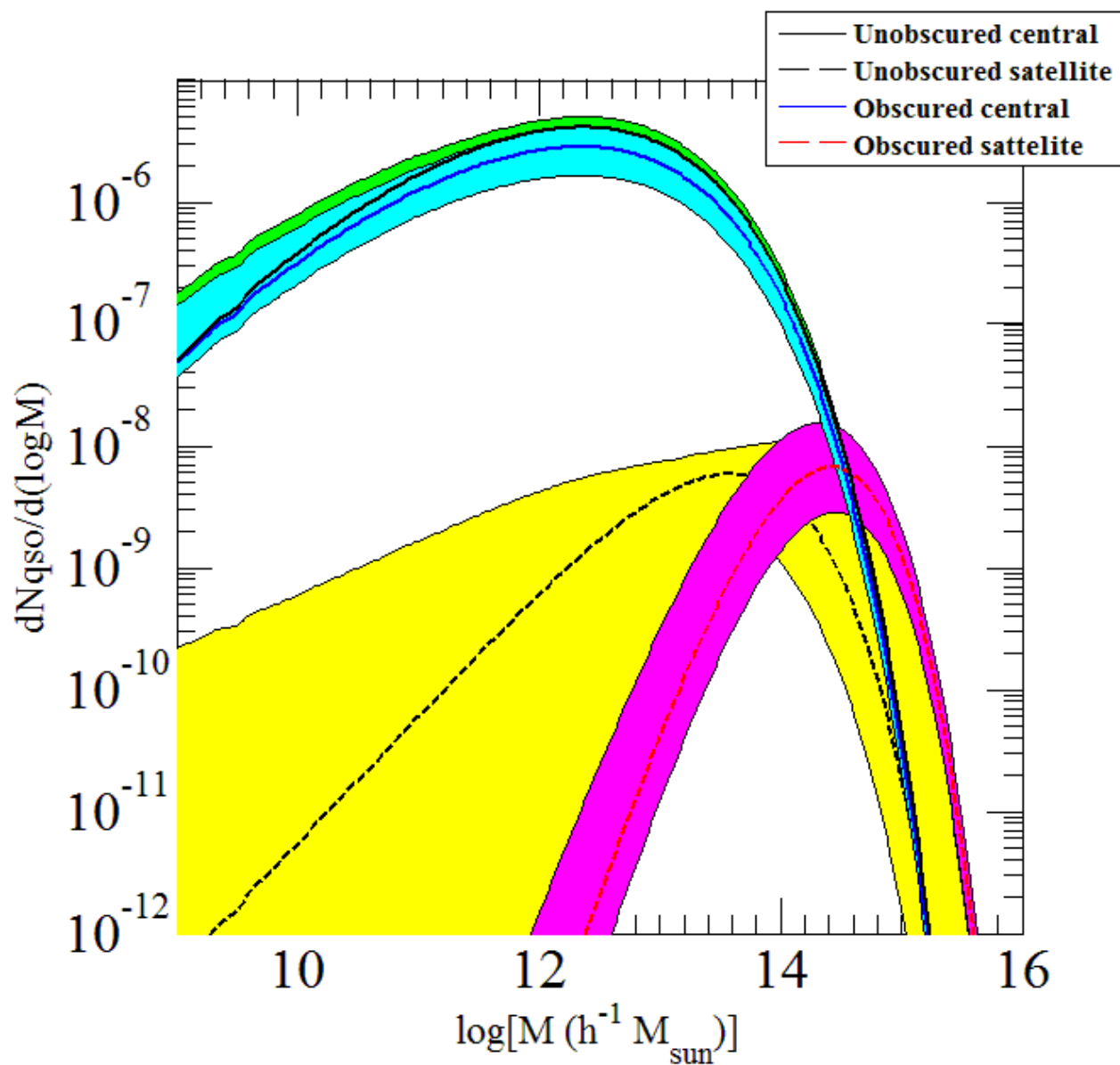
Obscured

Orientation
v/s
Evolutionary
Theories of
Quasars...?

Unobscured



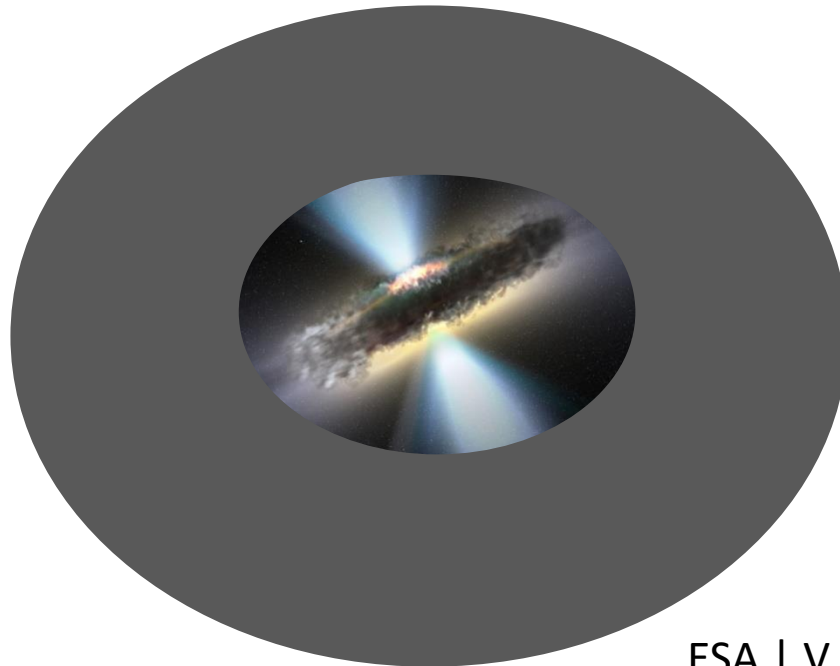
And Finally...



To quasars and beyond ...!!

- Our results match brilliantly with the Richardson et. al. 2012 results on HOD of SDSS quasars.
- The broader goal to compare with HOD of different classes of AGNs and analyze the galaxy-DM halo co-evolution.
- There are several scopes of improving, and we will do that, but as far as our preliminary results are concerned, the take-home-message is

Obscured and Unobscured Quasars are sitting in very similar Dark Matter Halos, and at least from the HOD perspective, the orientation theory of quasars strongly holds its ground.



Thank you !!!



S. Chatterjee



Z. Zheng



M. DiPompeo



R. Hickox



J. Richardson



A. Myers